

Physics 210B

L3 Beyond Boltzmann ... to Euler

So Far: - derived Boltzmann equation

$$\partial_t f + \underline{v} \cdot \underline{\nabla} f = C(f)$$

$$\left[\partial_t f + \underline{v} \cdot \underline{\nabla} f + \underline{g} \cdot \underline{\nabla}_v f = C(f) \right]$$

$$\downarrow$$

$$\frac{q}{m} \underline{E} + \frac{q}{mc} \underline{v} \times \underline{B} - \text{Vlasov for } C(f) \rightarrow$$

- proved H-Thm. $\left\{ \begin{array}{l} \text{Phase space} \\ \text{Fluid continuity} \end{array} \right.$

$$\frac{dS}{dt} \geq 0$$

→ identified Eqbm. Distribution f_{Max}

→ $f \rightarrow f_{\text{Max}}$ on τ_c time scale

→ local Maxwellian.

Now what?

What to do with Boltzmann Equation?

→ real problems:

- dynamics
 - inhomogeneity
- } combine collective and collisional processes

i.e. heat — conduction
 \ convection.

→ two roads forward:

- low collisionality, work with modified Vlasov equation

⇒ see Plasmas, Galactic Dynamics

$$d < n^{-1/3} < l_{mp} < L$$

- Fluid Equations

i.e. $f \sim f_{max} [T(\underline{x}), n(\underline{x}), \underline{V}(\underline{x})]$
 here

$\Rightarrow$ evolution equations for:
 $n(\underline{x}), \underline{V}(\underline{x}), T(\underline{x})$ etc.

$\Rightarrow$ macroscopic

$\Rightarrow$ Derived from Boltzmann...

Related: Inhomogeneity - i.e. $T(\underline{x})$

$\Rightarrow$ Transport

$$\underline{Q} = -\kappa \underline{\nabla} T \quad \left\{ \begin{array}{l} \text{constitutive} \\ \text{relation} \end{array} \right.$$

$\downarrow$
 thermal diffusivity
 $\rightarrow$ transport coefficient

$$\rightarrow f = f_{max.} + \delta f$$

i.e. distribution 'close to', but
 not equal to Maxwellian.
 $l_{mp}/L \ll 1$, but not $\rightarrow 0$

viscosity, heat conductivity..

- how compute transport coefficients.
⇒ needed for real fluid equations
- point:

$$\partial_t F + \underline{v} \cdot \underline{\nabla} F = C(F)$$

For inhomogeneous f_{eq} ,

$$\cancel{\partial_t f_{eq}} + \underline{v} \cdot \underline{\nabla} f_{eq} = C(\cancel{f_{eq}})$$

- does not satisfy Boltzmann equation!

$$- F = f_{eq} + \underset{\substack{\uparrow \\ \text{collisional Flux}}}{df}$$

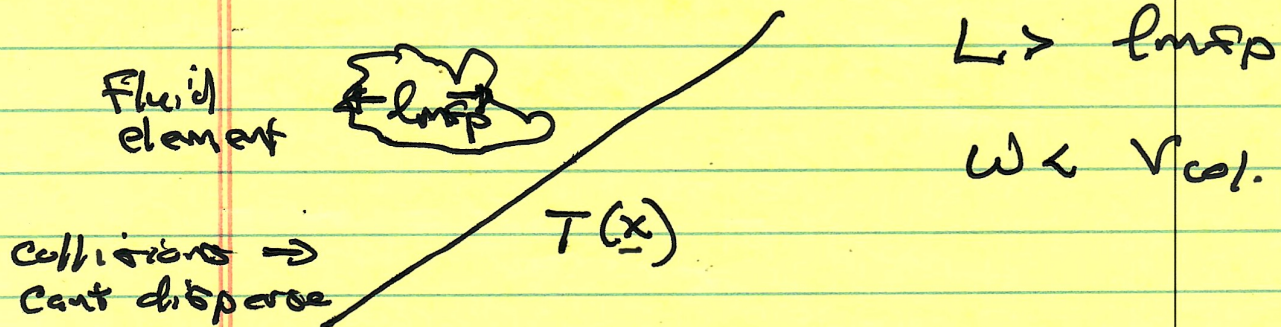
The Rest:

On Fluid Equations:

- Euler from Boltzmann ("ideal")
- Euler from macroscopics
- On Fluids (general thoughts),

Fluid Eqs

- replace B.E. by set of equations which evolve thermodynamic parameters.
- local Eulerian description (lab frame) i.e. $n(\underline{x}, t)$, $T(\underline{x}, t)$ etc
- describes 'blobs' of gas held together $\equiv$ by collisions.



- parametrized dynamics in terms

structure of distribution

$$f = \frac{n(\underline{x})}{(2\pi)^{3/2} v_{th}(\underline{x})^3} \exp \left[- \frac{(v - \underline{v}(\underline{x}, t))^2}{v_{th}(\underline{x}, t)^2} \right]$$

- works for slight deviation from equilibrium

i.e.

$$f = f_{eq} + \delta f$$

will see.

↓
local
Maxwellian

↓
Ideal Equations
(Perfect Fluid)

↓
Euler

= sacrifices higher moments of f .

$$\hookrightarrow \delta f \approx - \frac{\mathbf{v} \cdot \nabla}{v} f_{eq}$$

$$\approx \text{length } \nabla f_{eq}$$

$$\approx \boxed{\frac{\text{length } f_{eq}}{L}}$$

↓
viscous dissipative
equations

↓
Navier-Stokes

Ideal Equations:

$$\frac{\delta f}{\delta t} + \mathbf{v} \cdot \nabla f = C(f)$$

demand:

$$\int d^3v C(f) = 0$$

$$\int d^3v m v C(f) = 0$$

$$\int d^3v \mathbf{v} C(f) = 0$$

easily shown.
Collision operator
conserves
mass, energy
momentum.

So natural to define:

$$n = \int d^3v f \rightarrow \text{density}$$

$$\underline{V} = \underline{V}(\underline{x}, t) = \frac{1}{n} \int d^3v \underline{v} f \rightarrow \text{fluid / momentum.}$$

$$\bar{E} = \frac{1}{n} \int d^3v G f \rightarrow \text{energy density}$$

Now,

$$\partial_t f + \partial_{x_i} (V_i f) = C(f)$$

Conservative
Form

Taking moments:

$$\int d^3v *$$

$$\frac{\partial n}{\partial t} + \nabla \cdot (n \underline{V}) = 0$$

number/density
flux

cons.

Continuity,
cons.

$$\int d^3v \, m \, v_\alpha \neq$$

$$\Rightarrow$$

$$\partial_t (m \bar{v}_\alpha) + \frac{\partial}{\partial x_\beta} \pi_{\alpha, \beta} = 0$$

$\downarrow$
 mom. cons

$$\pi_{\alpha\beta} = \int d^3v \, m \, v_\alpha v_\beta F$$

$\downarrow$
 momentum flux

$$\text{and } \int d^3v \, \underline{v} \neq$$

$$\Rightarrow$$

$$\partial_t (\underline{v} \bar{E}) + \underline{\nabla} \cdot \underline{z} = 0$$

$$\underline{z} = \int d^3v \, \underline{v} \bar{E} F$$

Note: $\left. \begin{matrix} \underline{v} \\ \underline{E} \end{matrix} \right\}$ equations have the form:

①

$$\frac{d}{dt}(\text{stuff}) + \underline{\nabla} \cdot (\text{Flux of stuff}) = 0$$

i.e. of form:

$$\frac{\partial \rho}{\partial t} + \underline{\nabla} \cdot \underline{J} = 0$$

↖ macroscopic conservation

↓

② all rest upon conservation properties of the collision operator.

③ key is constitutive relation, i.e.

relating $\underline{J}$ to something useful.

i.e. Fick's Law: $\underline{J} = -D \underline{\nabla} \rho$

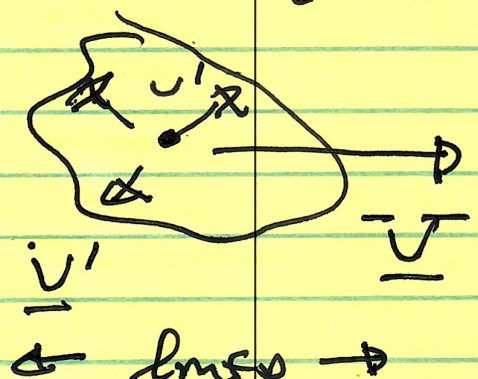
$$\Rightarrow \frac{\partial \rho}{\partial t} - D \nabla^2 \rho = 0$$

usually not so simple ...

→ closure problem.

④ Essence of fluid equation construction is calculation of fluxes

Further simplify:

$$\underset{\substack{\downarrow \\ \text{particle} \\ \text{velocity} \\ (\text{micro})}}{\underline{v}} = \underset{\substack{\downarrow \\ \text{Mean bulk} \\ \text{flow} \\ (\text{macro}) \\ \text{- linked to} \\ \text{body forces}}}{\underline{V}(\underline{x}, t)} + \underset{\substack{\text{thermal fluctuation} \\ \text{abt mean } (\sim \sqrt{T/m})}}{\underline{v}'}$$


Realistically, $|\underline{V}| < |\underline{v}'|$
but $\underline{v}'$'s cancel $\rightarrow$ random

or

$$\Pi_{\alpha\beta} = \int d^3v \quad m \left(\underline{v}_\alpha(\underline{x}, t) + \underline{v}'_\alpha \right) \left(\underline{v}_\beta(\underline{x}, t) + \underline{v}'_\beta \right) \rho$$

Now here,

$$F = F_{eZ} + \cancel{\sigma F}$$

$\Rightarrow$ ideal fluid

$\downarrow$ $\hookrightarrow$ would have σF dependence,
locally Maxwellian

and taking out $\underline{V}$, F_{eZ} is even in

σ

$$\Pi_{\alpha, \beta} = mn \left(V_{\alpha}(x, t) V_{\beta}(x, t) + \langle V'_{\alpha} V'_{\beta} \rangle \right)$$

$$F = \frac{n(x)}{(2\pi)^{3/2} [v_{th}(x)]^3} \exp \left[-\frac{V'^2}{v_{th}(x)^2} \right]$$

p.o. in $\frac{p_{MFP}}{L}$ $V' = V - \underline{V}(x, t)$

σ

$$\langle V'_{\alpha} V'_{\beta} \rangle = \frac{1}{3} V'^2 \delta_{\alpha, \beta}$$

(isotropic
 F_{eZ})

$$\langle V'^2 \rangle = 3T/m.$$

3x3

so, can define: $\underline{\underline{P}} = [\quad]$

$$\begin{aligned} \underline{\underline{P}} &= mn \langle v_\alpha' v_\beta' \rangle \rightarrow \text{stress tensor} \\ &= \frac{1}{3} mn \langle v'^2 \rangle \delta_{\alpha, \beta} \rightarrow \text{diagonal for ideal fluid} \end{aligned}$$

$\Rightarrow$ 1st moment:

$$\partial_t (n \underline{v}) + \nabla \cdot (n \underline{v} \underline{v} + \underline{\underline{P}}) = 0 \quad (*) \quad (\text{Euler})$$

↓
Reynolds stress tensor

$$\text{result: } \frac{\partial n}{\partial t} + \nabla \cdot (n \underline{v}) = 0$$

**

eliminate
and ~~subtract~~ ** from *:

Momentum
- $\nabla \cdot (n \underline{v} \underline{v})$ (Continuity)

$$n \left(\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} \right) = - \nabla P$$

" Euler
Eqn. "

→ microscopically, Euler Eqn.

corresponds to $\frac{\rho_{\text{max}}}{L} \rightarrow 0$.

→ $\underline{\underline{\underline{\Pi}}} = \rho \underline{\underline{\underline{V}}} \underline{\underline{\underline{V}}} + p \underline{\underline{\underline{I}}}$ is constitutive relation
for

$$\partial_t(\rho \underline{\underline{\underline{V}}}) + \underline{\underline{\underline{D}}} \cdot \underline{\underline{\underline{\Pi}}} = 0$$

→ Note C.R. contains no derivatives

ideal.

Contrast: Viscous fluid:

viscosity

$$\underline{\underline{\underline{\Pi}}} = \rho \underline{\underline{\underline{V}}} \underline{\underline{\underline{V}}} + p \underline{\underline{\underline{I}}} - \eta \underline{\underline{\underline{D}}} \underline{\underline{\underline{V}}}$$

↑
viscous stress

Viscous fluid

⇒ Navier Stokes

→ What is ρ ?

$\nabla \cdot \underline{Euler} = 0$
 for $\nabla \cdot \underline{v} = 0$

→ incompressible

$\nabla \cdot \underline{v} \neq 0 \Rightarrow$ Equation of state
i.e. $\rho = \rho(\dots)$

Similarly;

$$\begin{aligned}
 E &= \underbrace{\frac{1}{2} m v^2}_{KE} + \underbrace{E'}_{\text{internal dof}} \\
 &= \frac{1}{2} m (\underline{V}(\underline{x}, t) + \underline{v}')^2 + E'
 \end{aligned}$$

$$\begin{aligned}
 Z &= \int E \underline{v} f d^3v \\
 &\cong \int E \underline{v} \underbrace{f_{eq}}_{\uparrow} d^3v
 \end{aligned}$$

so

$$\underline{Z} = \int d^3v (\underline{V}(x,t) + \underline{v}) (E' + \frac{1}{2} (\underline{V}(x,t) + \underline{v})^2) f$$

$$F = f_{eq} \quad \text{PV work}$$

$$\underline{Z} = \underline{V}(x,t) \left(\frac{1}{2} n m \underline{V}^2 + \underbrace{P + n \bar{E}'}_{\text{enthalpy}} \right)$$

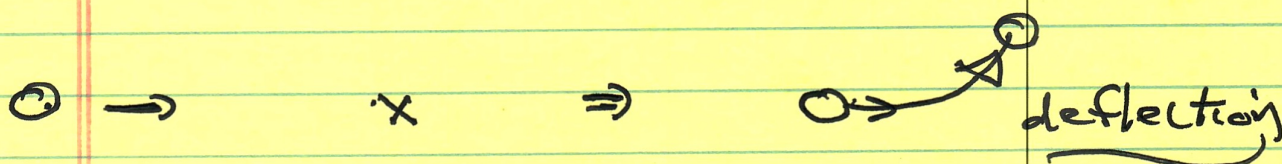
so

$$\underline{Z}_+ (n \bar{E}) + \underline{\nabla} \cdot \left[\underline{V}(x,t) \left(\frac{1}{2} n m \underline{V}^2 + P + n \bar{E}' \right) \right] = 0$$

can simplify, eq for momentum.

N.B:

angular momentum not conserved by CCF



→ Most truncations stop at 3rd moment.