

→ Collisional Regime

Here: $l_{co} < l_{mp} < l_c$

(short mean free path)

Point: → $l_{mp} < l_c$ ⇒ particle random walks parallel and undergoes many kicks in l_c . So parallel motion is diffusive.

→ perpendicular motion is continuous coarse graining/spreading, at $D_{\perp} \sim \rho_0^2 v_{th} \sim \rho_0^2 \frac{v_{th}}{l_{mp}}$

deriv.

So, can write:

$\langle dr^2 \rangle \sim D_{\parallel} l_{co} \sigma$

↓ parallel correlation length
(significant diffusive regime)

but also note that parallel motion is diffusive, so:

but time ~~set~~ ^{set} by

$$\kappa_{||} / l_{cd}^2 \sim 1/t$$

$$\Rightarrow \frac{\langle \sigma^2 \rangle}{t} \sim \frac{\kappa_{||}}{l_{cd}^2} D_M l_{cd} \quad \text{contrast} \quad \frac{CS}{l_{cd} \gg l_{msp}}$$

$$\sim D_M \frac{\kappa_{||}}{l_{cd}} \sim \boxed{D_M \kappa_{||} / l_{cd}}$$

$\sim D_M \text{vthe } \frac{D_{msp}}{l_{cd}} \ll \frac{\lambda_D}{\text{collisions}}$

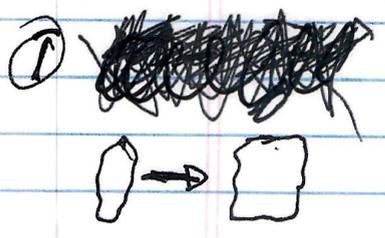
$$\boxed{\kappa_{\perp} = D_M \frac{\kappa_{||}}{l_{cd}}}$$

perpendicular heat conductivity in collisionless regime.

Now, what is l_{cd} ?

$$\begin{aligned} \kappa_{||} \kappa_{\perp} &\sim D_M^2 \frac{\partial^2}{\partial x^2} \\ \frac{\kappa_{||}}{l_c^2} &\sim \frac{1}{\delta^2} \\ l_{cd} &\sim \delta \left(\frac{\kappa_{||}}{D_M} \right)^{1/2} \end{aligned}$$

Notice l_{cd} is set by competition between 2 processes:



width δ increases due to diffusion (consequence)

$$\begin{aligned} l_{cd} &\sim \delta \left(\frac{\kappa_{||}}{D_M} \right)^{1/2} \\ \delta &\sim l_{cd} \left(\frac{D_M}{\kappa_{||}} \right)^{1/2} \end{aligned}$$

so $(d\sigma)^2 \sim (D_{\perp} dt)$
 $d\sigma \sim (D_{\perp} dt)^{1/2}$

but

$$\chi_{11} / (dL)^2 \approx 1/dt$$

⇒

$$d\sigma \sim \left(\frac{D_{\perp}}{\chi_{11}} (dL)^2 \right)^{1/2}$$

$$\boxed{d\sigma \sim \left(\frac{D_{\perp}}{\chi_{11}} \right)^{1/2} dL} \quad \checkmark$$

as before.

② width shrinks, due stochastic instability and area conservation: |



$$d\sigma/dL = -\sigma/\lambda_0$$

(exponential decay)

then balance at:

$$d\sigma \sim \left(D_{\perp} / \chi_{11} \right)^{1/2} dL \sim \underline{\sigma} dL$$

\downarrow
 smearing \leftarrow us \rightarrow thinning
 \swarrow \searrow

$\sigma \sim l_c \left(D_{\perp} / \chi_{11} \right)^{1/2}$

✓

M.B.: Can select σ from \downarrow

$$\partial_t T - \chi_{11} \nabla_{11}^2 T - D_{\perp} \nabla_{\perp}^2 T = 0$$

$$\Rightarrow \frac{\chi_{11}}{l_c^2} \sim \frac{D_{\perp}}{\sigma^2} \quad \underline{\sigma \sim l_c \left(D_{\perp} / \chi_{11} \right)^{1/2}}$$

Finally, need ~~correlation~~ length l_c for chunk size σ . Assume set by k_0



$$\underline{k_0^{-1}} \sim \underline{\sigma} e^{\frac{2}{3} / l_c} \Big|_{l_c} \approx \sigma e^{l_c / l_c}$$

$$\boxed{h_{cs} \sim h_c \ln(1/k_{cs})}$$

$$h_{cs} \approx h_c \ln \left(\frac{\kappa_{11}}{D_{\perp}} \right)^{1/2}$$

$$\boxed{h_{cs} \sim h_c \ln \left(\frac{\kappa_{11}}{D_{\perp}} \right)^{1/2} / k_{cs}}$$

$$\Rightarrow \boxed{\kappa_{\perp} \approx D_{\perp} \kappa_{11} / h_{cs}}$$

Apart from a log factor:

$$\kappa_{\perp} \approx v_{\perp} D_{\perp} \left(\frac{\kappa_{11}}{h_c} \right)$$

$\ll 1$

\Rightarrow reduced relative to collisionless values

Lesson: - collisions reduce (lump $\langle \epsilon \rangle$)
 reduce χ_{eff} relative to
 "Collisionless case"

- interplay of perp and parallel diffusion

- again, critical to knock particle off field line.

Now, the above calculation requires thought. Its much more convenient to crank ~~mindlessly~~ mindlessly.

Hydra approach: Kadomtsev and Pogutse (not mindless, but systematic)

Consider heat flux along wiggling fields
 d.e.

$$\underline{q} = -\chi_{||} \nabla_{||} T \hat{b} - \chi_{\perp} \nabla_{\perp} T$$

\downarrow parallel conduction \downarrow perp. conduction

Strickland codes

$$\chi_{||} \gg \chi_{\perp}$$

Here: $\underline{b} = \underline{b}_0 + \underline{\tilde{b}}$
 ↓ ↳ Fluctuations
 unperturbed

$\nabla_{11} = \partial_z + \underline{\tilde{b}} \cdot \underline{\nabla}_1$
 ↓
 piece along wiggling line

⚡, seeks mean radial heat flux

$\langle \dot{q}_r \rangle = -\kappa_{11} \langle b_r^2 \rangle \partial_r \langle T \rangle$ } usual quadratic
 $- \kappa_{11} \langle b_r \partial_z \tilde{T} \rangle$
 $- \kappa_{11} \langle b_r b_r \partial_r \tilde{T} \rangle$ → cubic
 $- \kappa_{11} \partial_r^2 \langle T \rangle$

Now $\underline{\textcircled{3}} \sim \frac{\kappa_{11} b_r b_r \tilde{T} / \Delta r}{\kappa_{11} b_r^2 / \Delta r}$

$\underline{\textcircled{2}} \sim \frac{\kappa_{11} b_r \tilde{T} / \Delta r}{\kappa_{11} b_r^2 / \Delta r}$

$\sim \frac{b_r \Delta r}{\Delta r} \sim \kappa_{11}$

so cubic nonlinearity dominates for $K_u > 1$.

$K_u < 1 \Rightarrow$ drop cubic.

To compute $\langle Z_r \rangle$, need

- retain ① (usual), and ②

- iterate for \tilde{T} using

$\nabla \cdot \tilde{z} = 0$ c.e. abt QLT.

Thinking (geop!) first:

$$\langle Z_r \rangle \approx -\psi_{11} \left[\langle k_{rr}^2 \rangle \partial_r T + \langle \tilde{b}_r \partial_z \tilde{T} \rangle - \psi_{\perp} \partial_r \langle T \rangle \right]$$

$$\approx -\psi_{11} \left[\langle \tilde{b}_r \overbrace{b \cdot \nabla T} \rangle - \psi_{\perp} \partial_r \langle T \rangle \right]$$

↙ identification:

$\partial_r \partial_r \langle T \rangle + \partial_z \tilde{T}$

Point: - need non-zero $\nabla \cdot \vec{T}$
fluctuation to drive heat flux
must have variation of T along line.

\rightarrow [c.e. temperature cant be constant along line, to drive parallel heat flux!]

- $\nabla \cdot \vec{q} = 0 \Rightarrow$ result must imply \underline{v}_\perp dependence!

B

$$\langle \vec{q}_\perp \rangle = -\nu_{||} \left[\langle \tilde{b}_r^2 \rangle \partial_r \langle T \rangle + \langle \tilde{b}_r \partial_z \tilde{T} \rangle \right]$$
$$\rightarrow \nu_\perp \nabla_\perp \langle T \rangle$$

$$\boxed{\nabla \cdot \vec{q} = 0}$$

$$\Rightarrow \nabla_{||} \tilde{q}_{||} + \nabla_\perp \cdot \tilde{\vec{q}}_\perp = -\nu_{||} \partial_z \tilde{b} \partial \langle T \rangle / \nu_r$$

c.e.

$$Q = -\chi_{||} \left[(\partial_z \pm \tilde{\omega} \cdot \underline{D}) (T_0 + \tilde{T}) (\underline{b}_0 + \tilde{\underline{b}}) \right] - \chi_{\perp} \underline{D}_{\perp} T$$

g11

$$- \chi_{||} \partial_z^2 \tilde{T} - \chi_{\perp} \underline{D}_{\perp}^2 \tilde{T} = -\chi_{||} \partial_z \tilde{\omega} \frac{\partial \langle T \rangle}{\partial r}$$

g11

$$\chi_{||} = \frac{-\chi_{||} \langle k_z \tilde{\omega}_{||} \rangle \partial \langle T \rangle / \partial r}{(\chi_{||} k_z^2 + \chi_{\perp} k_{\perp}^2)} \quad \text{for } T$$

g11

$$\begin{aligned} & \chi_{||} \langle \tilde{\omega}^2 \rangle \frac{\partial \langle T \rangle}{\partial r} - \chi_{||} \langle \tilde{\omega} \partial_z \tilde{T} \rangle \\ &= -\chi_{||} \sum_{||} \left(\frac{-\chi_{||} k_{||}^2 |\tilde{\omega}_{||}|^2}{\chi_{||} k_z^2 + \chi_{\perp} k_{\perp}^2} + |\tilde{\omega}_{\perp}|^2 \right) \frac{\partial \langle T \rangle}{\partial r} \\ &= -\chi_{||} \frac{\partial \langle T \rangle}{\partial r} \sum_{||} \left(\frac{-\chi_{||} k_{||}^2}{\chi_{||} k_z^2 + \chi_{\perp} k_{\perp}^2} + \frac{\chi_{||} k_{||}^2 + \chi_{\perp} k_{\perp}^2}{\chi_{||} k_z^2 + \chi_{\perp} k_{\perp}^2} \right) \end{aligned}$$

8/11

$$\langle Q_r \rangle_{NL} = -\chi_{11} \frac{\partial \langle T \rangle}{\partial r} \sum_n \frac{\chi_{\perp} k_{\perp}^2 \langle b_{n\perp}^2 \rangle}{\chi_{11} k_{n\parallel}^2 + \chi_{\perp} k_{\perp}^2}$$

Note explicit dependence on χ_{\perp} !

$$Z \sim (\chi_{11} v_{th})^{1/2} \sim D_B$$

80

$$\langle Q_r \rangle_{NL} \approx -\chi_{11} \frac{\partial \langle T \rangle}{\partial r} \int d\underline{k}_{\perp} \int d\underline{k}_{\parallel} \frac{\chi_{\perp} k_{\perp}^2 \langle \tilde{b}_{n\perp}^2 \rangle}{\chi_{11} (k_{\parallel}^2 + \frac{\chi_{\perp} k_{\perp}^2}{\chi_{11}})}$$

$$= -\frac{\partial \langle T \rangle}{\partial r} \int d\underline{k}_{\perp} \int d\underline{k}_{\parallel} \frac{\chi_{\perp} k_{\perp}^2 \langle \tilde{b}_{n\perp}^2 \rangle}{\left(\frac{k_{\parallel}^2}{(\chi_{\perp}/\chi_{11}) k_{\perp}^2} + 1 \right) \left(\frac{\chi_{\perp} k_{\perp}^2}{\chi_{11}} \right)}$$

$$= -\frac{\partial \langle T \rangle}{\partial r} \int d\underline{k}_{\perp} \frac{k_{\perp}^2 (\chi_{11} \chi_{\perp})^{1/2}}{\sqrt{k_{\perp}^2}} \langle \tilde{b}_{n\perp}^2 \rangle_{\text{acc}}$$

~~auto correlation~~
auto correlation

auto correlation ρ_{cc} enters via
normalization:

\Rightarrow

$$\langle q_r \rangle_{cc} \equiv -\sqrt{\kappa_{II} \chi_{\perp}} \langle \tilde{b}^2 \rangle_{cc} \sqrt{\kappa_{\perp}^2} \frac{dT}{dr}$$

Note: - need $\nabla_{II} \hat{T} \equiv -\overline{b_r} dT/dr$

($\underline{B} \cdot \underline{\nabla} T \neq 0$) for \perp heat flux

- $\langle \tilde{b}^2 \rangle_{cc} \sim D_M$

$\sqrt{\kappa_{\perp}^2} \sim 1/\Delta_{\perp}$

so

$$\langle q_r \rangle \equiv -\kappa_{I,eff} dT/dr - \chi_{\perp} dT/dr$$

$$\kappa_{I,eff} \equiv \sqrt{\kappa_{II} \chi_{\perp}} \frac{D_M}{\Delta_{\perp}}$$

$$\approx \frac{D_B D_M}{\Delta}$$

$$\left(\begin{array}{l} \chi_{\perp} \chi_{\perp} \sim \\ \frac{\chi_{\perp}^2}{\chi} \sim D_B \end{array} \right)$$

$$\chi_{\perp \text{eff}} \approx \frac{D_B D_M}{\Delta_{\perp}}$$

- $\chi_{\perp \text{eff}}$ scales with Bohm, not Spitzer (χ_{\parallel}) ✓

- kicking off line important, again.

$$\chi_{\perp} \sim D_M \frac{\chi_{\parallel}}{l_c} \sim v_{th} D_M \frac{l_{\text{res}}}{l_c}$$

To compare R & R:

$$\chi_{\perp} \approx \sqrt{\chi_{\parallel} \chi_{\perp}} \frac{\langle b^2 \rangle}{\Delta_{\perp}} l_{\text{res}}$$

what is Δ_{\perp} ?

Now

$$\frac{\chi_{\parallel}}{l_c^2} \sim \frac{\chi_{\perp}}{\Delta_{\perp}^2}$$

{ diffusion set
 \perp scale.

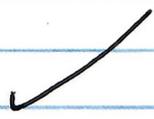
$$\Delta_{\perp} \sim l_c \sqrt{\chi_{\perp} / \chi_{\parallel}}$$

↑
 enters :
 spectrum,
 (small layer)

⇒

$$\chi_{\perp} \sim \frac{\sqrt{\chi_{\parallel} \chi_{\perp}} \langle \delta^2 \rangle l_c}{l_c (\chi_{\parallel} / \chi_{\perp})^{1/2}}$$

$$\chi_{\perp} \sim \frac{\chi_{\parallel} D_M}{l_c}$$



same as RR, to log.

so - modulo $k_{\perp}, \Lambda_{\perp}$; agrees with RR to within log. factor

$$\chi_{\perp} \sim v_{\perp} D_M \frac{l_{\text{mix}}}{l_c}$$

⇒ covers diffusion in $k_{\perp} \ll 1$ stochastic fields

⇒ Lesson: Take care re: irreversibility ↓